



## Introduction

- Dark matter remains to be one of the greatest mysteries of the universe
- Dark matter is thought to account for ~85% of the matter of the universe, yet has gone undetected due to its weak interaction with normal matter and no interaction with light
- Many contributions have been made to better understand the fundamental properties dark matter, yet no one has been able to detect it

## Evidence for Existence

- Since dark matter does not interact with light, the proof of the existence of dark matter comes from indirect observations

### Rotational Curves of Spiral Galaxies

- The relation between the velocity of stars in orbit around a galaxy and the distance between the stars and the center is given by the Keplerian curve
- According to this, the velocity of stars should fall at further distances from the center, but that's not what is found from observations (Sofue & Rubin, 2019)

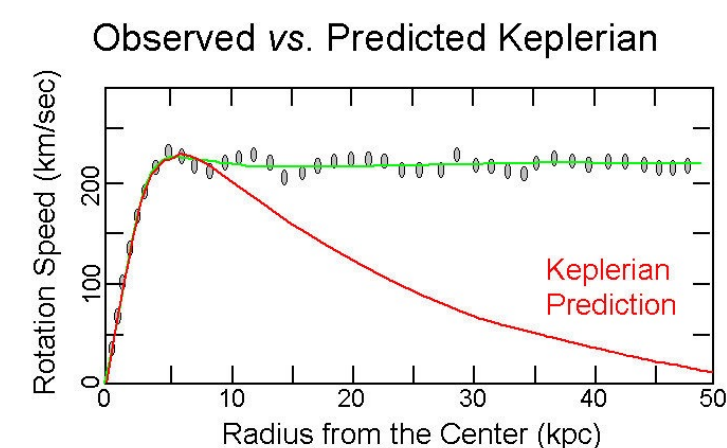


Fig 1: Predicted vs Observed Graphs of Rotational Spiral Galaxies

- The amount of luminous matter present is not enough for the stars near the edge to be moving at these high speeds (Bahcall, 2015)
- There must be other mass present, pointing to the possible existence of dark matter (Bahcall, 2015)

### Gravitational Lensing

- Gravitational lensing describes the phenomenon in which large amounts of matter creates a gravitational field that distorts and magnifies light coming from distant galaxies (Massey et al., 2010)
- Gravitational lensing provides information in the foreground based on its effects on the background galaxies (Massey et al., 2010)
- The amount of distortion of light provides strong evidence for the existence of vast amounts of dark matter within and in between galaxies (Bahcall, 2015)

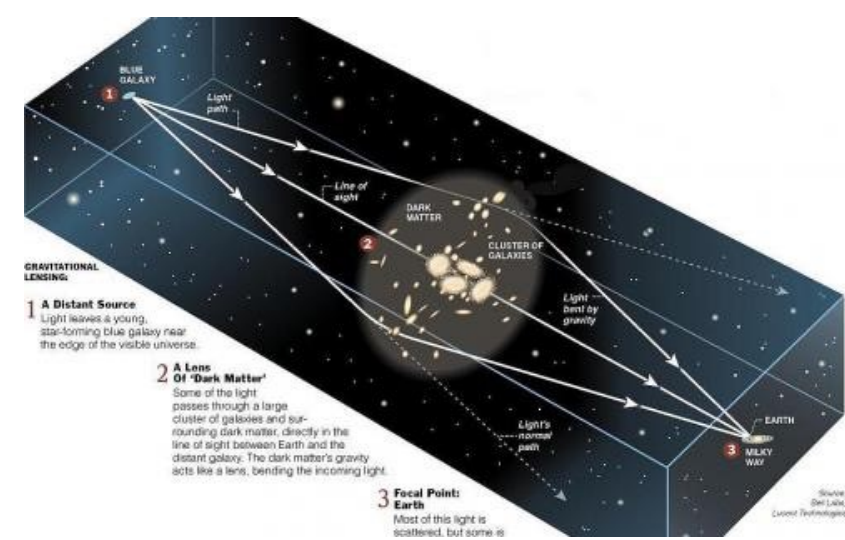


Fig 2: Gravitational Lensing Diagram

## Cosmic Microwave Background Radiation

- The cosmic microwave background radiation shows what the universe looked like before galaxies and clusters were formed (Jones & Lasenby, 1998)
- Ordinary matter interacts strongly with radiation while dark matter does not
- Ordinary and dark matter perturbations leave different imprints on the CMB, again pointing to the existence of dark matter (Bahcall, 2015)

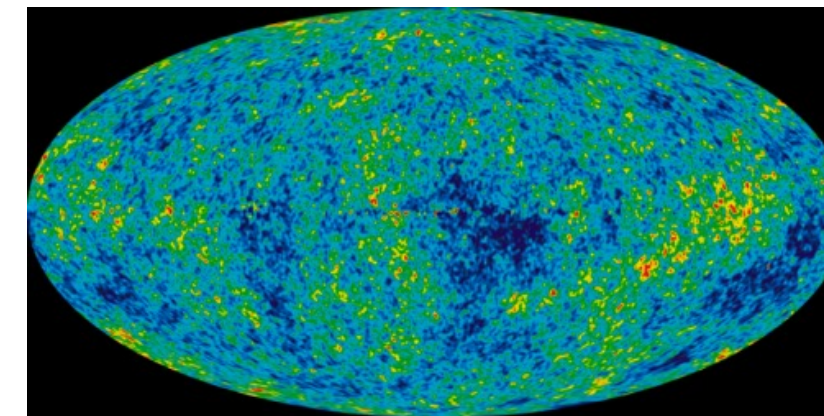


Fig 3: Cosmic Microwave Background Radiation Diagram

## Current Experiment Methods

- Although the existence of dark matter is widely accepted by physicists around the world, the form and properties of dark matter have not yet been determined
- There are three ways in which physicist attempt to detect dark matter

### Make Dark Matter in Accelerators

- One methods of searching for dark matter is to generate them by smashing protons together at nearly the speed of light (Chodos, 2004)
- Violent head-on collisions convert energy into showers of exotic particles scattering in all directions (Lowette, 2016)
- The hope is that within this debris, a short-lived dark matter particle will blink into existence (Lowette, 2016)
- This is not something new. The same method was used over a decade ago to discover the Higgs boson
- It is being used again by the European Organization for Nuclear Research, or CERN to discover the existence of dark matter

### Indirect Detection of Dark Matter

- This method of searching for dark matter focuses on looking for the products of dark matter interactions rather than the particle themselves (Leane, 2020)
- In general, one looks for gamma-rays, cosmic-rays, or neutrinos
- For example, although dark matter may not interact with light, the debris from two colliding dark matter particles may produce things such as gamma-rays, which are detectable (Leane, 2020)
- Some instruments that have been used to detect dark matter indirectly include H.E.S.S., VERITAS, and MAGIC

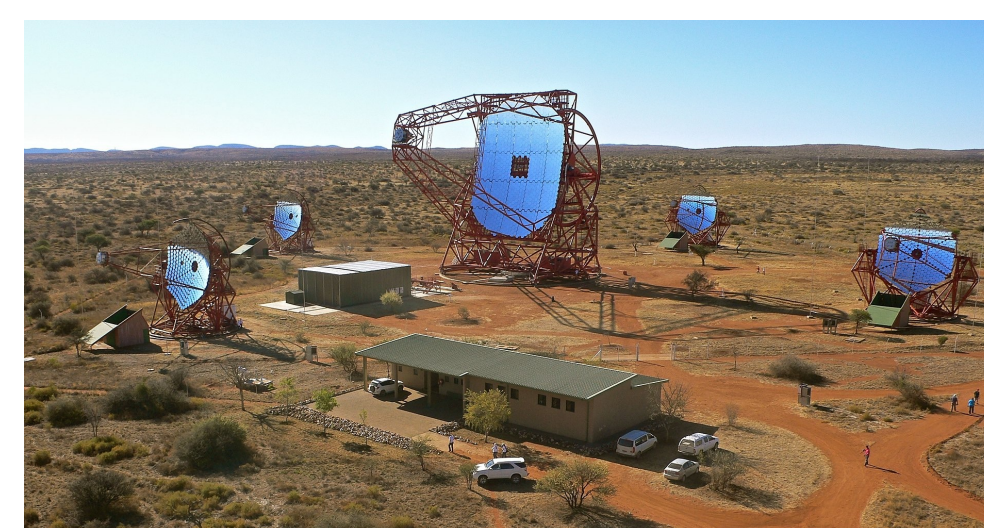


Fig 4: H.E.S.S. Telescope

## Direct Detection of Dark Matter

- Direct detection of dark matter attempts to directly measure dark matter collisions on Earth (Billard et al., 2021)
- A big problem is that the surface of the Earth is bombarded by cosmic rays, which create too much noise to be able to use direct detection
- Therefore, detectors designed to directly detect dark matter are placed far underground and inside mountains, where there is barely any noise (Billard et al., 2021)
- One example of an experiment using direct detection is XENON in Italy at the Gran Sasso Laboratory
- It is the brainchild of physicist Elena Aprile at Columbia University
- The experiment uses an instrument called the dual-phase, xenon-based time-projection chamber
- The very first iteration of the XENON series was XENON10 in 2007
- Competing experiments include LUX, ZEPLIN, and PANDAX



Fig 5: Dual-Phase, Xenon-Based Time-Projection Chamber

## Results and Conclusion

- There is a great amount of indirect evidence pointing to the existence of dark matter
- No experiment has been successful in uncovering the mystery of dark matter
- Physicists continue to develop better experiments in order to finally detect dark matter
- Understanding the properties of dark matter is important because it is thought to have played a fundamental role in the formation of galaxies and can be the answer to many observations

## Results and Conclusion

- Bahcall, N. A. (2015). Dark matter universe, PNAS **112**, 12243; DOI: 10.1073/pnas.1516944112
- Sofue, Y., & Rubin, V. (2019). Rotation curves of spiral galaxies. Annu. Rev. Astron. Astrophys. **39**, 137; DOI: 10.1146/annurev.astro.39.1.137
- Massey, R., Kitching, T., & Richard, J. (2010). The dark matter of gravitational lensing. Rep. Prog. Phys. **73**, 086901; DOI: 10.1088/0034-4885/73/8/086901
- Jones, A. W., & Lasenby, A. N. (1998). The cosmic microwave background. Living Rev. Relativ. **1**, 11; DOI: 10.12942/lrr-1998-11
- Chodos, A. (2004). Next-generation accelerators could hold key to dark matter, energy. APS News **13**, 7; https://www.aps.org/publications/apsnews/200407/accelerators.cfm
- Lowette, S. (2016). Accelerator searches for new physics in the context of dark matter. J. Phys.: Conf. Ser. **718**, 022011; DOI: 10.1088/1742-6596/718/2/022011
- Leane, R. K. (2020). Indirect detection of dark matter in the galaxy. arXiv:2006.00513; DOI: 10.48550/arXiv.2006.00513
- Billard, J., Boulay, M., Cebrián, S., Covi, L., Fiorillo, G., Green, A., Kopp, J., Majorovits, B., Palladino, K., Petricca, F., Roszkowski, L., & Schumann, M. (2022). Direct detection of dark matter -- APPEC committee report. Rep. Prog. Phys. **85**, 056201; DOI: 10.1088/1361-6633/ac5754